

# New Techniques for the $n$ -Body Problem

Andreas Knauf

Department Mathematik  
Universität Erlangen-Nürnberg (Germany)

Hamiltonian systems and their applications

Euler Institute (St. Petersburg)     June 3, 2015

# Overview

- Measure of the collision orbits
- Measure of non-collision singularities (2 bodies,  $n$  centers)
- Regularization of the Kepler problem
- Regularization of (simultaneous) binary collisions)
- Kinematics
- Train correspondences
- Classical scattering
- Open questions

with Stefan Fleischer, respectively with  
Jacques Fejoz and Richard Montgomery

Unpublished!

# Overview

- Measure of the collision orbits
- Measure of non-collision singularities (2 bodies,  $n$  centers)
- Regularization of the Kepler problem
- Regularization of (simultaneous) binary collisions)
- Kinematics
- Train correspondences
- Classical scattering
- Open questions

with Stefan Fleischer, respectively with  
Jacques Fejoz and Richard Montgomery

Unpublished!

# Overview

- Measure of the collision orbits
- Measure of non-collision singularities (2 bodies,  $n$  centers)
- Regularization of the Kepler problem
  - Regularization of (simultaneous) binary collisions)
  - Kinematics
  - Train correspondences
  - Classical scattering
  - Open questions

with Stefan Fleischer, respectively with Jacques Fejoz and Richard Montgomery

Unpublished!

# Overview

- Measure of the collision orbits
- Measure of non-collision singularities (2 bodies,  $n$  centers)
- Regularization of the Kepler problem
- Regularization of (simultaneous) binary collisions)
- Kinematics
- Train correspondences
- Classical scattering
- Open questions

with Stefan Fleischer, respectively with  
Jacques Fejoz and Richard Montgomery

Unpublished!

# Overview

- Measure of the collision orbits
- Measure of non-collision singularities (2 bodies,  $n$  centers)
- Regularization of the Kepler problem
- Regularization of (simultaneous) binary collisions)
- Kinematics
- Train correspondences
- Classical scattering
- Open questions

with Stefan Fleischer, respectively with  
Jacques Fejoz and Richard Montgomery

Unpublished!

# Overview

- Measure of the collision orbits
- Measure of non-collision singularities (2 bodies,  $n$  centers)
- Regularization of the Kepler problem
- Regularization of (simultaneous) binary collisions)
- Kinematics
- Train correspondences
- Classical scattering
- Open questions

with Stefan Fleischer, respectively with  
Jacques Fejoz and Richard Montgomery

Unpublished!

# Overview

- Measure of the collision orbits
- Measure of non-collision singularities (2 bodies,  $n$  centers)
- Regularization of the Kepler problem
- Regularization of (simultaneous) binary collisions)
- Kinematics
- Train correspondences
- Classical scattering
- Open questions

with Stefan Fleischer, respectively with  
Jacques Fejoz and Richard Montgomery

Unpublished!



# Overview

- Measure of the collision orbits
- Measure of non-collision singularities (2 bodies,  $n$  centers)
- Regularization of the Kepler problem
- Regularization of (simultaneous) binary collisions)
- Kinematics
- Train correspondences
- Classical scattering
- Open questions

with Stefan Fleischer, respectively with  
Jacques Fejoz and Richard Montgomery

Unpublished!

# Overview

- Measure of the collision orbits
- Measure of non-collision singularities (2 bodies,  $n$  centers)
- Regularization of the Kepler problem
- Regularization of (simultaneous) binary collisions)
- Kinematics
- Train correspondences
- Classical scattering
- Open questions

with Stefan Fleischer, respectively with  
Jacques Fejoz and Richard Montgomery

Unpublished!

# Overview

- Measure of the collision orbits
- Measure of non-collision singularities (2 bodies,  $n$  centers)
- Regularization of the Kepler problem
- Regularization of (simultaneous) binary collisions)
- Kinematics
- Train correspondences
- Classical scattering
- Open questions

with Stefan Fleischer, respectively with Jacques Fejoz and Richard Montgomery

Unpublished!

# Notation for $n$ bodies

## Case of celestial mechanics

- For  $n \in \mathbb{N}$  particles in
- $d \in \mathbb{N}$  degrees of freedom
- the configuration space is  $\widehat{M} := \mathbb{R}^{nd} \setminus \Delta$ ,
- with collision set  
 $\Delta := \{q = (q_1, \dots, q_n) \in \mathbb{R}^{nd} \mid q_i = q_j \text{ for some } i \neq j\}$ .
- Given masses  $m_1, \dots, m_n > 0$  and
- interaction parameters  $l_{i,j} \in \mathbb{R} \setminus \{0\}$

we consider the Hamiltonian

$$\widehat{H} : T^*\widehat{M} \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) := K(p) + V(q), \quad (1)$$

$$\text{with } K(p) := \sum_{i=1}^n \frac{\|p_i\|^2}{2m_i} \quad \text{and} \quad V(q) := \sum_{1 \leq i < j \leq n} \frac{l_{i,j}}{\|q_i - q_j\|}.$$

# Notation for $n$ bodies

## Case of celestial mechanics

- For  $n \in \mathbb{N}$  particles in
- $d \in \mathbb{N}$  degrees of freedom
- the configuration space is  $\widehat{M} := \mathbb{R}^{nd} \setminus \Delta$ ,
- with collision set  
 $\Delta := \{q = (q_1, \dots, q_n) \in \mathbb{R}^{nd} \mid q_i = q_j \text{ for some } i \neq j\}$ .
- Given masses  $m_1, \dots, m_n > 0$  and
- interaction parameters  $l_{i,j} \in \mathbb{R} \setminus \{0\}$

we consider the Hamiltonian

$$\widehat{H} : T^*\widehat{M} \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) := K(p) + V(q), \quad (1)$$

$$\text{with } K(p) := \sum_{i=1}^n \frac{\|p_i\|^2}{2m_i} \quad \text{and} \quad V(q) := \sum_{1 \leq i < j \leq n} \frac{l_{i,j}}{\|q_i - q_j\|}.$$

# Notation for $n$ bodies

## Case of celestial mechanics

- For  $n \in \mathbb{N}$  particles in
- $d \in \mathbb{N}$  degrees of freedom
- the configuration space is  $\hat{M} := \mathbb{R}^{nd} \setminus \Delta$ ,
- with collision set  
 $\Delta := \{q = (q_1, \dots, q_n) \in \mathbb{R}^{nd} \mid q_i = q_j \text{ for some } i \neq j\}$ .
- Given masses  $m_1, \dots, m_n > 0$  and
- interaction parameters  $l_{i,j} \in \mathbb{R} \setminus \{0\}$

we consider the Hamiltonian

$$\hat{H} : T^*\hat{M} \rightarrow \mathbb{R} \quad , \quad \hat{H}(p, q) := K(p) + V(q), \quad (1)$$

$$\text{with } K(p) := \sum_{i=1}^n \frac{\|p_i\|^2}{2m_i} \quad \text{and} \quad V(q) := \sum_{1 \leq i < j \leq n} \frac{l_{i,j}}{\|q_i - q_j\|}.$$

# Notation for $n$ bodies

## Case of celestial mechanics

- For  $n \in \mathbb{N}$  particles in
- $d \in \mathbb{N}$  degrees of freedom
- the configuration space is  $\widehat{M} := \mathbb{R}^{nd} \setminus \Delta$ ,
- with collision set  
 $\Delta := \{q = (q_1, \dots, q_n) \in \mathbb{R}^{nd} \mid q_i = q_j \text{ for some } i \neq j\}$ .
- Given masses  $m_1, \dots, m_n > 0$  and
- interaction parameters  $l_{i,j} \in \mathbb{R} \setminus \{0\}$

we consider the Hamiltonian

$$\widehat{H} : T^*\widehat{M} \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) := K(p) + V(q), \quad (1)$$

$$\text{with } K(p) := \sum_{i=1}^n \frac{\|p_i\|^2}{2m_i} \quad \text{and} \quad V(q) := \sum_{1 \leq i < j \leq n} \frac{l_{i,j}}{\|q_i - q_j\|}.$$

# Notation for $n$ bodies

## Case of celestial mechanics

- For  $n \in \mathbb{N}$  particles in
- $d \in \mathbb{N}$  degrees of freedom
- the configuration space is  $\widehat{M} := \mathbb{R}^{nd} \setminus \Delta$ ,
- with collision set  
 $\Delta := \{q = (q_1, \dots, q_n) \in \mathbb{R}^{nd} \mid q_i = q_j \text{ for some } i \neq j\}$ .
- Given masses  $m_1, \dots, m_n > 0$  and
- interaction parameters  $l_{i,j} \in \mathbb{R} \setminus \{0\}$

we consider the Hamiltonian

$$\widehat{H} : T^*\widehat{M} \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) := K(p) + V(q), \quad (1)$$

$$\text{with } K(p) := \sum_{i=1}^n \frac{\|p_i\|^2}{2m_i} \quad \text{and} \quad V(q) := \sum_{1 \leq i < j \leq n} \frac{l_{i,j}}{\|q_i - q_j\|}.$$



# Notation for $n$ bodies

## Case of celestial mechanics

- For  $n \in \mathbb{N}$  particles in
- $d \in \mathbb{N}$  degrees of freedom
- the configuration space is  $\widehat{M} := \mathbb{R}^{nd} \setminus \Delta$ ,
- with collision set  
 $\Delta := \{q = (q_1, \dots, q_n) \in \mathbb{R}^{nd} \mid q_i = q_j \text{ for some } i \neq j\}$ .
- Given masses  $m_1, \dots, m_n > 0$  and
- interaction parameters  $l_{i,j} \in \mathbb{R} \setminus \{0\}$

we consider the Hamiltonian

$$\widehat{H} : T^*\widehat{M} \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) := K(p) + V(q), \quad (1)$$

$$\text{with } K(p) := \sum_{i=1}^n \frac{\|p_i\|^2}{2m_i} \quad \text{and} \quad V(q) := \sum_{1 \leq i < j \leq n} \frac{l_{i,j}}{\|q_i - q_j\|}.$$

# Notation for $n$ bodies

## Case of celestial mechanics

- For  $n \in \mathbb{N}$  particles in
- $d \in \mathbb{N}$  degrees of freedom
- the configuration space is  $\widehat{M} := \mathbb{R}^{nd} \setminus \Delta$ ,
- with collision set  
 $\Delta := \{q = (q_1, \dots, q_n) \in \mathbb{R}^{nd} \mid q_i = q_j \text{ for some } i \neq j\}$ .
- Given masses  $m_1, \dots, m_n > 0$  and
- interaction parameters  $l_{i,j} \in \mathbb{R} \setminus \{0\}$

we consider the Hamiltonian

$$\widehat{H} : T^*\widehat{M} \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) := K(p) + V(q), \quad (1)$$

$$\text{with } K(p) := \sum_{i=1}^n \frac{\|p_i\|^2}{2m_i} \quad \text{and} \quad V(q) := \sum_{1 \leq i < j \leq n} \frac{l_{i,j}}{\|q_i - q_j\|}.$$

# Notation for $n$ bodies

## Case of celestial mechanics

- For  $n \in \mathbb{N}$  particles in
- $d \in \mathbb{N}$  degrees of freedom
- the configuration space is  $\widehat{M} := \mathbb{R}^{nd} \setminus \Delta$ ,
- with collision set  
 $\Delta := \{q = (q_1, \dots, q_n) \in \mathbb{R}^{nd} \mid q_i = q_j \text{ for some } i \neq j\}$ .
- Given masses  $m_1, \dots, m_n > 0$  and
- interaction parameters  $l_{i,j} \in \mathbb{R} \setminus \{0\}$

we consider the Hamiltonian

$$\widehat{H} : T^*\widehat{M} \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) := K(p) + V(q), \quad (1)$$

$$\text{with } K(p) := \sum_{i=1}^n \frac{\|p_i\|^2}{2m_i} \quad \text{and} \quad V(q) := \sum_{1 \leq i < j \leq n} \frac{l_{i,j}}{\|q_i - q_j\|}.$$

# Notions for $n$ body motion

- Phase space  $\hat{P} := T^*\hat{M}$ ,
- Maximal flow  $\Phi = (p, q) : D \rightarrow \hat{P}$  with domain

$$D = \{(t, x) \in \mathbb{R} \times \hat{P} \mid T^-(x) < t < T^+(x)\}.$$

## Definition

- Set  $\text{Sing} := \{x \in \hat{P} \mid T^+(x) < +\infty\}$  of phase space points encountering a **singularity** in the future;
- Subset of phase space points encountering a **collision** singularity:

$$\text{Coll} := \left\{ x \in \text{Sing} \mid \lim_{t \nearrow T^+(x)} q(t, x) \text{ exists} \right\}.$$

# Measure of collision orbits

## Theorem (Saari)

For dimension  $d = 3$  and forces  $\nabla V_{i,j}(q) = -m_i m_j q / \|q\|^{1+\alpha}$

- collisions are improbable for  $\alpha < 17/7$ .
- Binary collisions are improbable for  $\alpha < 3$ .

Don Saari: Improbability of Collisions in Newtonian Gravitational Systems. II. *Transactions of the AMS* **181**, 351–368 (1973)

**Generalisation:** Dimension  $d \geq 2$

$$V(q) = \sum_{1 \leq i < j \leq n} V_{i,j}(q_i - q_j) \quad \text{with} \quad V_{i,j} \in C^2(\mathbb{R}^d \setminus \{0\}, \mathbb{R})$$

## Definition

A potential  $V$  is called

- *moderate* if  $\|\nabla V_{i,j}(q)\| \leq C\|q\|^{-3+\varepsilon}$  for  $\|q\| \leq 1$ ,
- *long ranged* if  $\|\nabla V_{i,j}(q)\| \leq C\|q\|^{-1-\varepsilon}$  for  $\|q\| \geq 1$ .

# Measure of collision orbits

## Theorem (Saari)

For dimension  $d = 3$  and forces  $\nabla V_{i,j}(q) = -m_i m_j q / \|q\|^{1+\alpha}$

- collisions are improbable for  $\alpha < 17/7$ .
- Binary collisions are improbable for  $\alpha < 3$ .

Don Saari: Improbability of Collisions in Newtonian Gravitational Systems. II. *Transactions of the AMS* **181**, 351–368 (1973)

**Generalisation:** Dimension  $d \geq 2$

$$V(q) = \sum_{1 \leq i < j \leq n} V_{i,j}(q_i - q_j) \quad \text{with} \quad V_{i,j} \in C^2(\mathbb{R}^d \setminus \{0\}, \mathbb{R})$$

## Definition

A potential  $V$  is called

- *moderate* if  $\|\nabla V_{i,j}(q)\| \leq C\|q\|^{-3+\varepsilon}$  for  $\|q\| \leq 1$ ,
- *long ranged* if  $\|\nabla V_{i,j}(q)\| \leq C\|q\|^{-1-\varepsilon}$  for  $\|q\| \geq 1$ .

# Measure of collision orbits

## Theorem (St. Fleischer, AK)

*For moderate (and long ranged) potentials  $V$  collisions are improbable.*

**Proof.** Use family of global Poincaré surfaces of section  $\mathcal{H}^{(m)}$  in phase space 'around' the collision set  $\Delta$ , with

- surface area  $\int_{\mathcal{H}^{(m)}} i_{X_H} \Omega$  (with phase space volume form  $\Omega$ ) going to zero as  $m \rightarrow \infty$ ,
- and any collision orbit intersecting all  $\mathcal{H}^{(m)}$  with  $m$  large.
- Cluster: set partition  $\mathcal{C} = \{C_1, \dots, C_k\}$  of  $\{1, \dots, n\}$
- $\text{Coll}_{\mathcal{C}} := \{x \in \text{Coll} \mid \lim_{t \nearrow T^+(x)} (q_i(t, x) - q_j(t, x)) = 0 \text{ iff } i \cong_{\mathcal{C}} j\}$ ,

$$\text{Coll} = \bigcup_{\mathcal{C}} \text{Coll}_{\mathcal{C}}.$$

- Use internal and external coordinates for cluster  $\mathcal{C}$  to define  $\mathcal{H}^{(m)}$ :  
modulus of internal position as function of other coordinates.



# Measure of collision orbits

## Theorem (St. Fleischer, AK)

*For moderate (and long ranged) potentials  $V$  collisions are improbable.*

**Proof.** Use family of global Poincaré surfaces of section  $\mathcal{H}^{(m)}$  in phase space 'around' the collision set  $\Delta$ , with

- surface area  $\int_{\mathcal{H}^{(m)}} \mathbf{i}_{X_H} \Omega$  (with phase space volume form  $\Omega$ ) going to zero as  $m \rightarrow \infty$ ,
- and any collision orbit intersecting all  $\mathcal{H}^{(m)}$  with  $m$  large.
- **Cluster:** set partition  $\mathcal{C} = \{C_1, \dots, C_k\}$  of  $\{1, \dots, n\}$
- $\text{Coll}_{\mathcal{C}} := \{x \in \text{Coll} \mid \lim_{t \nearrow T^+(x)} (q_i(t, x) - q_j(t, x)) = 0 \text{ iff } i \cong_{\mathcal{C}} j\}$ ,

$$\text{Coll} = \bigcup_{\mathcal{C}} \text{Coll}_{\mathcal{C}}.$$

- Use internal and external coordinates for cluster  $\mathcal{C}$  to define  $\mathcal{H}^{(m)}$ : modulus of internal position as function of other coordinates.





# Measure of collision orbits

## Theorem (St. Fleischer, AK)

*For moderate (and long ranged) potentials  $V$  collisions are improbable.*

**Proof.** Use family of global Poincaré surfaces of section  $\mathcal{H}^{(m)}$  in phase space 'around' the collision set  $\Delta$ , with

- surface area  $\int_{\mathcal{H}^{(m)}} \mathbf{i}_{X_H} \Omega$  (with phase space volume form  $\Omega$ ) going to zero as  $m \rightarrow \infty$ ,
- and any collision orbit intersecting all  $\mathcal{H}^{(m)}$  with  $m$  large.
- Cluster: set partition  $\mathcal{C} = \{C_1, \dots, C_k\}$  of  $\{1, \dots, n\}$
- $\text{Coll}_{\mathcal{C}} := \{x \in \text{Coll} \mid \lim_{t \nearrow T^+(x)} (q_i(t, x) - q_j(t, x)) = 0 \text{ iff } i \cong_{\mathcal{C}} j\}$ ,

$$\text{Coll} = \bigcup_{\mathcal{C}} \text{Coll}_{\mathcal{C}}.$$

- Use internal and external coordinates for cluster  $\mathcal{C}$  to define  $\mathcal{H}^{(m)}$ :  
modulus of internal position as function of other coordinates.



# Measure of collision orbits

## Theorem (St. Fleischer, AK)

*For moderate (and long ranged) potentials  $V$  collisions are improbable.*

**Proof.** Use family of global Poincaré surfaces of section  $\mathcal{H}^{(m)}$  in phase space 'around' the collision set  $\Delta$ , with

- surface area  $\int_{\mathcal{H}^{(m)}} \mathbf{i}_{X_H} \Omega$  (with phase space volume form  $\Omega$ ) going to zero as  $m \rightarrow \infty$ ,
- and any collision orbit intersecting all  $\mathcal{H}^{(m)}$  with  $m$  large.
- **Cluster:** set partition  $\mathcal{C} = \{C_1, \dots, C_k\}$  of  $\{1, \dots, n\}$
- $\text{Coll}_{\mathcal{C}} := \{x \in \text{Coll} \mid \lim_{t \nearrow T^+(x)} (q_i(t, x) - q_j(t, x)) = 0 \text{ iff } i \cong_{\mathcal{C}} j\}$ ,

$$\text{Coll} = \bigcup_{\mathcal{C}} \text{Coll}_{\mathcal{C}}.$$

- Use internal and external coordinates for cluster  $\mathcal{C}$  to define  $\mathcal{H}^{(m)}$ :  
modulus of internal position as function of other coordinates.



# Measure of collision orbits

## Theorem (St. Fleischer, AK)

*For moderate (and long ranged) potentials  $V$  collisions are improbable.*

**Proof.** Use family of global Poincaré surfaces of section  $\mathcal{H}^{(m)}$  in phase space 'around' the collision set  $\Delta$ , with

- surface area  $\int_{\mathcal{H}^{(m)}} \mathbf{i}_{X_H} \Omega$  (with phase space volume form  $\Omega$ ) going to zero as  $m \rightarrow \infty$ ,
- and any collision orbit intersecting all  $\mathcal{H}^{(m)}$  with  $m$  large.
- **Cluster:** set partition  $\mathcal{C} = \{C_1, \dots, C_k\}$  of  $\{1, \dots, n\}$
- **Coll $_{\mathcal{C}}$**  :=  $\{x \in \text{Coll} \mid \lim_{t \nearrow T+(x)} (q_i(t, x) - q_j(t, x)) = 0 \text{ iff } i \cong_{\mathcal{C}} j\}$ ,

$$\text{Coll} = \dot{\bigcup}_{\mathcal{C}} \text{Coll}_{\mathcal{C}}.$$

- Use internal and external coordinates for cluster  $\mathcal{C}$  to define  $\mathcal{H}^{(m)}$ :  
modulus of internal position as function of other coordinates.



# Measure of collision orbits

## Theorem (St. Fleischer, AK)

*For moderate (and long ranged) potentials  $V$  collisions are improbable.*

**Proof.** Use family of global Poincaré surfaces of section  $\mathcal{H}^{(m)}$  in phase space 'around' the collision set  $\Delta$ , with

- surface area  $\int_{\mathcal{H}^{(m)}} \mathbf{i}_{X_H} \Omega$  (with phase space volume form  $\Omega$ ) going to zero as  $m \rightarrow \infty$ ,
- and any collision orbit intersecting all  $\mathcal{H}^{(m)}$  with  $m$  large.
- **Cluster:** set partition  $\mathcal{C} = \{C_1, \dots, C_k\}$  of  $\{1, \dots, n\}$
- **Coll $_{\mathcal{C}}$**  :=  $\{x \in \text{Coll} \mid \lim_{t \nearrow T^+(x)} (q_i(t, x) - q_j(t, x)) = 0 \text{ iff } i \cong_{\mathcal{C}} j\}$ ,

$$\text{Coll} = \dot{\bigcup}_{\mathcal{C}} \text{Coll}_{\mathcal{C}}.$$

- Use internal and external coordinates for cluster  $\mathcal{C}$  to define  $\mathcal{H}^{(m)}$ : modulus of internal position as function of other coordinates.

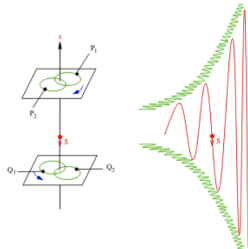


# Measure of non-collision singularities

For an  $n$ -center – two-body system

- Non-collision singularities exist in the four-body problem.

Z. Xia: The existence of noncollision singularities in Newtonian systems. *Annals of Mathematics* **135**, 411–468 (1992)



- They are improbable in the four-body problem.
- Non-collision singularities exist in a simplified four-body problem.

D. Saari: Improbability of Collisions in Newtonian Gravitational Systems. *Transactions AMS* **162**, 267–271 (1971)

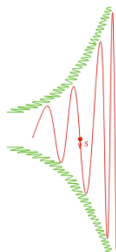
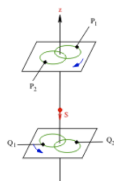
J. Xue, D. Dolgopyat: Non-collision singularities in the planar two-center-two-body problem. arXiv:1307.2645, 2013

# Measure of non-collision singularities

For an  $n$ -center – two-body system

- Non-collision singularities exist in the four-body problem.

Z. Xia: The existence of noncollision singularities in Newtonian systems. *Annals of Mathematics* **135**, 411–468 (1992)



- They are improbable in the four-body problem.

D. Saari: Improbability of Collisions in Newtonian Gravitational Systems. *Transactions AMS* **162**, 267–271 (1971)

- Non-collision singularities exist in a simplified four-body problem.

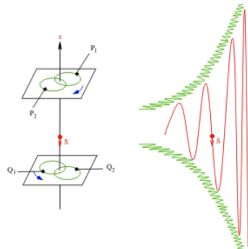
J. Xue, D. Dolgopyat: Non-collision singularities in the planar two-center-two-body problem. arXiv:1307.2645, 2013

# Measure of non-collision singularities

For an  $n$ -center – two-body system

- Non-collision singularities exist in the four-body problem.

Z. Xia: The existence of noncollision singularities in Newtonian systems. *Annals of Mathematics* **135**, 411–468 (1992)



- They are improbable in the four-body problem.

D. Saari: Improbability of Collisions in Newtonian Gravitational Systems. *Transactions AMS* **162**, 267–271 (1971)

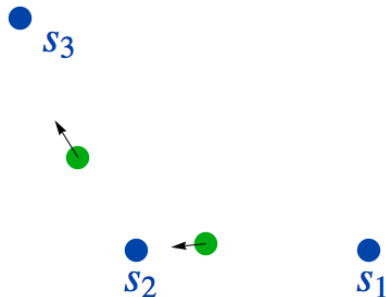
- Non-collision singularities exist in a simplified four-body problem.

J. Xue, D. Dolgopyat: Non-collision singularities in the planar two-center-two-body problem. arXiv:1307.2645, 2013

# Measure of non-collision singularities

Theorem (St. Fleischer, AK)

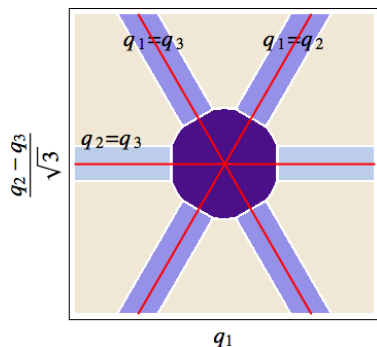
*For moderate pair potentials non-collision singularities are improbable.*



- Fixed centers  $s_1, \dots, s_n \in \mathbb{R}^d$
- two moving particles



# Poincaré surfaces for collision set



Picture in configuration space, for

- $d = 1$  dimension,
- $n = 3$  bodies,
- in center of mass system.

Collision set  $\Delta$

## Why another regularisation?

There are already regularisations by  
Levi-Civita, Kustaanheimo-Stiefel, Moser, Souriau...

Advantages:

- All dimensions  $d \geq 1$
- simultaneously for all energies
- no change of time parameter
- symplectic extension of phase space.

## Why another regularisation?

There are already regularisations by  
Levi-Civita, Kustaanheimo-Stiefel, Moser, Souriau...

Advantages:

- All dimensions  $d \geq 1$
- simultaneously for all energies
- no change of time parameter
- symplectic extension of phase space.

## Why another regularisation?

There are already regularisations by  
Levi-Civita, Kustaanheimo-Stiefel, Moser, Souriau...

Advantages:

- All dimensions  $d \geq 1$
- simultaneously for all energies
- no change of time parameter
- symplectic extension of phase space.

## Why another regularisation?

There are already regularisations by  
Levi-Civita, Kustaanheimo-Stiefel, Moser, Souriau...

Advantages:

- All dimensions  $d \geq 1$
- simultaneously for all energies
- no change of time parameter
- symplectic extension of phase space.

## Why another regularisation?

There are already regularisations by  
Levi-Civita, Kustaanheimo-Stiefel, Moser, Souriau...

Advantages:

- All dimensions  $d \geq 1$
- simultaneously for all energies
- no change of time parameter
- symplectic extension of phase space.

## Why another regularisation?

There are already regularisations by  
Levi-Civita, Kustaanheimo-Stiefel, Moser, Souriau...

Advantages:

- All dimensions  $d \geq 1$
- simultaneously for all energies
- no change of time parameter
- symplectic extension of phase space.

## What kind of regularity?

- regularity of the single collision orbits, e.g. analyticity in  $t^{1/3}$
- regularity of a Poincaré map (block regularity)
- regularity of the flow.



## What **kind** of regularity?

- regularity of the single collision orbits, e.g. analyticity in  $t^{1/3}$
- regularity of a Poincaré map (block regularity)
- regularity of the flow.

## What **kind** of regularity?

- regularity of the single collision orbits, *e.g.* analyticity in  $t^{1/3}$
- regularity of a Poincaré map (block regularity)
- regularity of the flow.

## What kind of regularity?

- regularity of the single collision orbits, *e.g.* analyticity in  $t^{1/3}$
- regularity of a Poincaré map (block regularity)
- regularity of the flow.

# Regularisation by phase space extension: Kepler case

- Hamiltonian, attractive case ( $I < 0$ ):

$$\widehat{H} : T^*(\mathbb{R}^d \setminus \{0\}) \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) = \frac{\|p\|^2}{2m} + \frac{I}{\|q\|}$$

- $\varepsilon$ -neighborhood of the excluded fiber  $T_0^*\mathbb{R}^d$

$$\widehat{U}^\varepsilon := \left\{ (p, q) \in T^*(\mathbb{R}^d \setminus \{0\}) \mid \|q\| < \varepsilon, \frac{\|p\|^2}{2m} > \frac{3}{4} \frac{|I|}{\|q\|} \right\}$$

- $\widehat{T} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}$ : the time elapsed since the closest encounter of the Kepler solution with the pericentre.
- Laplace-Runge-Lenz vector (with  $\widehat{L}(p, q) := q \wedge p$ )

$$\widetilde{A} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d \quad , \quad \widetilde{A}(p, q) := -\widehat{L}(p, q)p + mI \frac{q}{\|q\|}.$$

- its direction  $\widehat{A} : \widehat{U}^\varepsilon \rightarrow S^{d-1}$  ,  $\widehat{A} := \widetilde{A} / \|\widetilde{A}\|$  .
- $\widehat{B} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d$  ,  $\widehat{B} := \widehat{L}\widehat{A}$

# Regularisation by phase space extension: Kepler case

- Hamiltonian, attractive case ( $I < 0$ ):

$$\widehat{H} : T^*(\mathbb{R}^d \setminus \{0\}) \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) = \frac{\|p\|^2}{2m} + \frac{I}{\|q\|}$$

- $\varepsilon$ -neighborhood of the excluded fiber  $T_0^*\mathbb{R}^d$

$$\widehat{U}^\varepsilon := \left\{ (p, q) \in T^*(\mathbb{R}^d \setminus \{0\}) \mid \|q\| < \varepsilon, \frac{\|p\|^2}{2m} > \frac{3}{4} \frac{|I|}{\|q\|} \right\}$$

- $\widehat{T} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}$ : the time elapsed since the closest encounter of the Kepler solution with the pericentre.
- Laplace-Runge-Lenz vector (with  $\widehat{L}(p, q) := q \wedge p$ )

$$\widetilde{A} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d \quad , \quad \widetilde{A}(p, q) := -\widehat{L}(p, q)p + mI \frac{q}{\|q\|}.$$

- its direction  $\widehat{A} : \widehat{U}^\varepsilon \rightarrow S^{d-1}$  ,  $\widehat{A} := \widetilde{A} / \|\widetilde{A}\|$  .
- $\widehat{B} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d$  ,  $\widehat{B} := \widehat{L}\widehat{A}$

# Regularisation by phase space extension: Kepler case

- Hamiltonian, attractive case ( $I < 0$ ):

$$\widehat{H} : T^*(\mathbb{R}^d \setminus \{0\}) \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) = \frac{\|p\|^2}{2m} + \frac{I}{\|q\|}$$

- $\varepsilon$ -neighborhood of the excluded fiber  $T_0^*\mathbb{R}^d$

$$\widehat{U}^\varepsilon := \left\{ (p, q) \in T^*(\mathbb{R}^d \setminus \{0\}) \mid \|q\| < \varepsilon, \frac{\|p\|^2}{2m} > \frac{3}{4} \frac{|I|}{\|q\|} \right\}$$

- $\widehat{T} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}$ : the time elapsed since the closest encounter of the Kepler solution with the pericentre.
- Laplace-Runge-Lenz vector (with  $\widehat{L}(p, q) := q \wedge p$ )

$$\widetilde{A} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d \quad , \quad \widetilde{A}(p, q) := -\widehat{L}(p, q)p + ml \frac{q}{\|q\|}.$$

- its direction  $\widehat{A} : \widehat{U}^\varepsilon \rightarrow S^{d-1}$  ,  $\widehat{A} := \widetilde{A} / \|\widetilde{A}\|$  .
- $\widehat{B} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d$  ,  $\widehat{B} := \widehat{L}\widehat{A}$

# Regularisation by phase space extension: Kepler case

- Hamiltonian, attractive case ( $I < 0$ ):

$$\widehat{H} : T^*(\mathbb{R}^d \setminus \{0\}) \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) = \frac{\|p\|^2}{2m} + \frac{I}{\|q\|}$$

- $\varepsilon$ -neighborhood of the excluded fiber  $T_0^*\mathbb{R}^d$

$$\widehat{U}^\varepsilon := \left\{ (p, q) \in T^*(\mathbb{R}^d \setminus \{0\}) \mid \|q\| < \varepsilon, \frac{\|p\|^2}{2m} > \frac{3}{4} \frac{|I|}{\|q\|} \right\}$$

- $\widehat{T} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}$ : the time elapsed since the closest encounter of the Kepler solution with the pericentre.
- Laplace-Runge-Lenz vector (with  $\widehat{L}(p, q) := q \wedge p$ )

$$\widetilde{A} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d \quad , \quad \widetilde{A}(p, q) := -\widehat{L}(p, q)p + mI \frac{q}{\|q\|}.$$

- its direction  $\widehat{A} : \widehat{U}^\varepsilon \rightarrow S^{d-1}$  ,  $\widehat{A} := \widetilde{A} / \|\widetilde{A}\|$  .
- $\widehat{B} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d$  ,  $\widehat{B} := \widehat{L}\widehat{A}$

# Regularisation by phase space extension: Kepler case

- Hamiltonian, attractive case ( $I < 0$ ):

$$\widehat{H} : T^*(\mathbb{R}^d \setminus \{0\}) \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) = \frac{\|p\|^2}{2m} + \frac{I}{\|q\|}$$

- $\varepsilon$ -neighborhood of the excluded fiber  $T_0^*\mathbb{R}^d$

$$\widehat{U}^\varepsilon := \left\{ (p, q) \in T^*(\mathbb{R}^d \setminus \{0\}) \mid \|q\| < \varepsilon, \frac{\|p\|^2}{2m} > \frac{3}{4} \frac{|I|}{\|q\|} \right\}$$

- $\widehat{T} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}$ : the time elapsed since the closest encounter of the Kepler solution with the pericentre.
- Laplace-Runge-Lenz vector (with  $\widehat{L}(p, q) := q \wedge p$ )

$$\widetilde{A} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d \quad , \quad \widetilde{A}(p, q) := -\widehat{L}(p, q)p + mI \frac{q}{\|q\|}.$$

- its direction  $\widehat{A} : \widehat{U}^\varepsilon \rightarrow S^{d-1}$  ,  $\widehat{A} := \widetilde{A}/\|\widetilde{A}\|$  .
- $\widehat{B} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d$  ,  $\widehat{B} := \widehat{L}\widehat{A}$



# Regularisation by phase space extension: Kepler case

- Hamiltonian, attractive case ( $I < 0$ ):

$$\widehat{H} : T^*(\mathbb{R}^d \setminus \{0\}) \rightarrow \mathbb{R} \quad , \quad \widehat{H}(p, q) = \frac{\|p\|^2}{2m} + \frac{I}{\|q\|}$$

- $\varepsilon$ -neighborhood of the excluded fiber  $T_0^*\mathbb{R}^d$

$$\widehat{U}^\varepsilon := \left\{ (p, q) \in T^*(\mathbb{R}^d \setminus \{0\}) \mid \|q\| < \varepsilon, \frac{\|p\|^2}{2m} > \frac{3}{4} \frac{|I|}{\|q\|} \right\}$$

- $\widehat{T} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}$ : the time elapsed since the closest encounter of the Kepler solution with the pericentre.
- Laplace-Runge-Lenz vector (with  $\widehat{L}(p, q) := q \wedge p$ )

$$\widetilde{A} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d \quad , \quad \widetilde{A}(p, q) := -\widehat{L}(p, q)p + mI \frac{q}{\|q\|}.$$

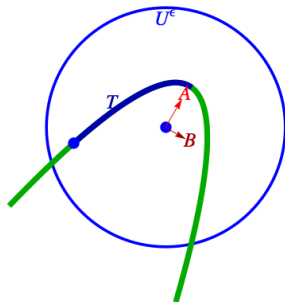
- its direction  $\widehat{A} : \widehat{U}^\varepsilon \rightarrow S^{d-1}$  ,  $\widehat{A} := \widetilde{A}/\|\widetilde{A}\|$  .
- $\widehat{B} : \widehat{U}^\varepsilon \rightarrow \mathbb{R}^d$  ,  $\widehat{B} := \widehat{L}\widehat{A}$

# Regularisation: Kepler case

Together we obtain a smooth map

$$\widehat{\Psi} := (\widehat{H}, \widehat{T}, \widehat{A}, \widehat{B}) : \widehat{U}^\varepsilon \longrightarrow T^*(\mathbb{R} \times S^{d-1}),$$

- all entries except time  $\widehat{T}$  are constants of the motion.
- $\text{im}(\widehat{\Psi})$  misses the zero section  $\cong \mathbb{R} \times S^{d-1}$  of the cotangent bundle, since the collision orbits are characterized by zero angular momentum and thus  $\widehat{B} = \widehat{L}\widehat{A} = 0$ , and the point of collision on such an orbit corresponds to  $\widehat{T} \rightarrow 0$ .

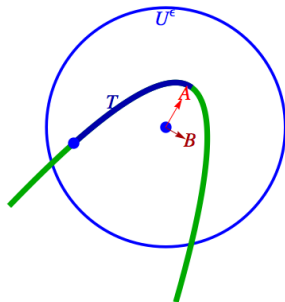


# Regularisation: Kepler case

Together we obtain a smooth map

$$\widehat{\Psi} := (\widehat{H}, \widehat{T}, \widehat{A}, \widehat{B}) : \widehat{U}^\varepsilon \longrightarrow T^*(\mathbb{R} \times S^{d-1}),$$

- all entries except time  $\widehat{T}$  are constants of the motion.
- $\text{im}(\widehat{\Psi})$  misses the zero section  $\cong \mathbb{R} \times S^{d-1}$  of the cotangent bundle, since the collision orbits are characterized by zero angular momentum and thus  $\widehat{B} = \widehat{L}\widehat{A} = 0$ , and the point of collision on such an orbit corresponds to  $\widehat{T} \rightarrow 0$ .

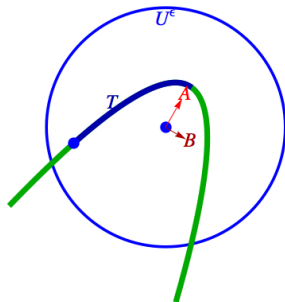


# Regularisation: Kepler case

Together we obtain a smooth map

$$\widehat{\Psi} := (\widehat{H}, \widehat{T}, \widehat{A}, \widehat{B}) : \widehat{U}^\varepsilon \longrightarrow T^*(\mathbb{R} \times S^{d-1}),$$

- all entries except time  $\widehat{T}$  are constants of the motion.
- $\text{im}(\widehat{\Psi})$  misses the zero section  $\cong \mathbb{R} \times S^{d-1}$  of the cotangent bundle, since the collision orbits are characterized by zero angular momentum and thus  $\widehat{B} = \widehat{L}\widehat{A} = 0$ , and the point of collision on such an orbit corresponds to  $\widehat{T} \rightarrow 0$ .



# Regularisation: Kepler case

- We now complete phase space by setting

$$P := T^*(\mathbb{R}^d \setminus \{0\}) \cup (\mathbb{R} \times S^{d-1})$$

and introducing a second chart  $\Psi$  with domain

$$U^\varepsilon := \widehat{U}^\varepsilon \cup (\mathbb{R} \times S^{d-1}) \subseteq P,$$

$$\Psi \equiv (H, T, A, B) : U^\varepsilon \rightarrow T^*(\mathbb{R} \times S^{d-1}),$$

$$\Psi|_{\widehat{U}^\varepsilon} := \widehat{\Psi} \quad \text{and} \quad \Psi|_{\mathbb{R} \times S^{d-1}}(h, a) := (h, 0, a, 0).$$

## Lemma

$\widehat{\Psi}$  is a diffeomorphism onto its image. On  $U^\varepsilon$  for all  $1 \leq i, j \leq d$

$$\{H, T\} = 1 \quad , \quad \{A_i, B_j\} = \delta_{i,j} - A_i A_j \quad , \quad \{B_i, B_j\} = L_{i,j},$$

all other Poisson brackets being zero.

The Hamiltonian system  $(P, \omega, H)$  is a *smooth complete* extension of  $(T^*(\mathbb{R}^d \setminus \{0\}), dq \wedge dp, \widehat{H})$ .

# $n$ -body simultaneous binary regularisation:

non-smooth dynamics

- McGehee (1974): Triple collisions are not regularizable.
  - So regularize all (simultaneous) binary collisions by extending  $T^*\widehat{M}$  (with  $\widehat{M} = \mathbb{R}^{nd} \setminus \Delta$ ) to a smooth manifold  $P$  like above.
  - Then orbits without higher order near-collision exist for all times.
  - Martínez–Simó: Poincaré map no smoother than  $C^{8/3}$ .
- Regularisation (ElBialy): Simultaneous binary collisions are continuous (and become rest points).

R. Martínez, C. Simó: The degree of differentiability of the regularization of simultaneous binary collisions in some  $N$ -body problems. *Nonlinearity* 13, 2107–2130 (2000)

M. ElBialy: Collective branch regularization of simultaneous binary collisions in the 3D  $N$ -body problem. *J. Math. Phys.* 50 052702 (2009)

# $n$ -body simultaneous binary regularisation:

non-smooth dynamics

- McGehee (1974): **Triple** collisions are not regularizable.
- So regularize all (simultaneous) **binary** collisions by extending  $T^*\widehat{M}$  (with  $\widehat{M} = \mathbb{R}^{nd} \setminus \Delta$ ) to a smooth manifold  $P$  like above.
- Then orbits without higher order near-collision exist for all times.
- Martínez–Simó: Poincaré map no smoother than  $C^{8/3}$ .  
R. Martínez, C. Simó: The degree of differentiability of the regularization of simultaneous binary collisions in some  $N$ -body problems. *Nonlinearity* **13**, 2107–2130 (2000)
- Regularisation (ElBialy): Simultaneous binary collisions are **continuous** (and become rest points).  
M. ElBialy: Collective branch regularization of simultaneous binary collisions in the 3D  $N$ -body problem. *J. Math. Phys.* **50** 052702 (2009)



# $n$ -body simultaneous binary regularisation:

non-smooth dynamics

- McGehee (1974): Triple collisions are not regularizable.
- So regularize all (simultaneous) binary collisions by extending  $T^*\widehat{M}$  (with  $\widehat{M} = \mathbb{R}^{nd} \setminus \Delta$ ) to a smooth manifold  $P$  like above.
- Then orbits without higher order near-collision exist for all times.
- Martínez–Simó: Poincaré map no smoother than  $C^{8/3}$ .
- Regularisation (ElBialy): Simultaneous binary collisions are continuous (and become rest points).

R. Martínez, C. Simó: The degree of differentiability of the regularization of simultaneous binary collisions in some  $N$ -body problems. *Nonlinearity* 13, 2107–2130 (2000)

M. ElBialy: Collective branch regularization of simultaneous binary collisions in the 3D  $N$ -body problem. *J. Math. Phys.* 50 052702 (2009)

# $n$ -body simultaneous binary regularisation:

non-smooth dynamics

- McGehee (1974): Triple collisions are not regularizable.
- So regularize all (simultaneous) binary collisions by extending  $T^*\hat{M}$  (with  $\hat{M} = \mathbb{R}^{nd} \setminus \Delta$ ) to a smooth manifold  $P$  like above.
- Then orbits without higher order near-collision exist for all times.
- Martínez–Simó: Poincaré map no smoother than  $C^{8/3}$ .
- Regularisation (ElBialy): Simultaneous binary collisions are continuous (and become rest points).

R. Martínez, C. Simó: The degree of differentiability of the regularization of simultaneous binary collisions in some  $N$ -body problems. *Nonlinearity* 13, 2107–2130 (2000)

M. ElBialy: Collective branch regularization of simultaneous binary collisions in the 3D  $N$ -body problem. *J. Math. Phys.* 50 052702 (2009)

# $n$ -body simultaneous binary regularisation:

non-smooth dynamics

- McGehee (1974): Triple collisions are not regularizable.
- So regularize all (simultaneous) binary collisions by extending  $T^*\hat{M}$  (with  $\hat{M} = \mathbb{R}^{nd} \setminus \Delta$ ) to a smooth manifold  $P$  like above.
- Then orbits without higher order near-collision exist for all times.
- Martínez–Simó: Poincaré map no smoother than  $C^{8/3}$ .
- Regularisation (ElBialy): Simultaneous binary collisions are continuous (and become rest points).

R. Martínez, C. Simó: The degree of differentiability of the regularization of simultaneous binary collisions in some  $N$ -body problems. *Nonlinearity* **13**, 2107–2130 (2000)

M. ElBialy: Collective branch regularization of simultaneous binary collisions in the 3D  $N$ -body problem. *J. Math. Phys.* **50** 052702 (2009)

# $n$ -body simultaneous binary regularisation:

non-smooth dynamics

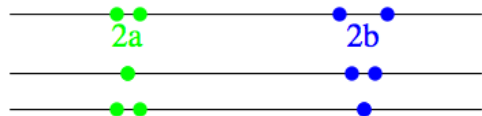
- McGehee (1974): Triple collisions are not regularizable.
- So regularize all (simultaneous) binary collisions by extending  $T^*\widehat{M}$  (with  $\widehat{M} = \mathbb{R}^{nd} \setminus \Delta$ ) to a smooth manifold  $P$  like above.
- Then orbits without higher order near-collision exist for all times.
- Martínez–Simó: Poincaré map no smoother than  $C^{8/3}$ .  
R. Martínez, C. Simó: The degree of differentiability of the regularization of simultaneous binary collisions in some  $N$ -body problems. *Nonlinearity* **13**, 2107–2130 (2000)
- Regularisation (ElBialy): Simultaneous binary collisions are continuous (and become rest points).  
M. ElBialy: Collective branch regularization of simultaneous binary collisions in the 3D  $N$ -body problem. *J. Math. Phys.* **50** 052702 (2009)

# Intuition for the Martínez–Simó $C^{8/3}$ result

This is for special one-parameter solutions, the parameter  $\Delta t \in \mathbb{R}$  being the time between the two binary collisions.

**Example:** Four equal masses on a line, distance one at simultaneous binary collision.

Distance between first pair:  $2a(t)$ , between second pair:  $2b(t)$ .



Force on second pair:  $\ddot{b} = -\frac{1}{2}b^{-2} +$

$$\frac{1}{2} \left( (1+a-b)^{-2} + (1-a-b)^{-2} - (1+a+b)^{-2} - (1-a+b)^{-2} \right) \\ = -\frac{1}{2}b^{-2} + 4b + 8b^3 + 24a^2b + \dots$$

Perturbation theory:  $b(t) = \left(\frac{3}{2}t\right)^{2/3} + \dots$ ,  $a(t) = \left(\frac{3}{2}(t + \Delta t)\right)^{2/3} + \dots$

# Intuition for the Martínez–Simó $C^{8/3}$ result

- Rate of change in energy of  $b$  pair, for  $t$  of same order as  $\Delta t$ :

$$\ddot{b}b = \dots + k a^2 \dot{b}b = \dots + k \operatorname{sign}(\Delta t) \Delta t^{5/3}.$$

- Integrate for time  $\sim \Delta t$ :  $k \operatorname{sign}(\Delta t) \Delta t^{8/3}$
- This  $C^{8/3}$  result is valid for the celestial mechanics case.  
Less differentiability for (attracting) charged particles.

# Intuition for the Martínez–Simó $C^{8/3}$ result

- Rate of change in energy of  $b$  pair, for  $t$  of same order as  $\Delta t$ :

$$\ddot{b}b = \dots + k a^2 \dot{b}b = \dots + k \operatorname{sign}(\Delta t) \Delta t^{5/3}.$$

- Integrate for time  $\sim \Delta t$ :  $k \operatorname{sign}(\Delta t) \Delta t^{8/3}$
- This  $C^{8/3}$  result is valid for the celestial mechanics case.  
Less differentiability for (attracting) charged particles.

# Intuition for the Martínez–Simó $C^{8/3}$ result

- Rate of change in energy of  $b$  pair, for  $t$  of same order as  $\Delta t$ :  
 $\ddot{b}b = \dots + k a^2 \dot{b}b = \dots + k \operatorname{sign}(\Delta t) \Delta t^{5/3}$ .
- Integrate for time  $\sim \Delta t$ :  $k \operatorname{sign}(\Delta t) \Delta t^{8/3}$
- This  $C^{8/3}$  result is valid for the celestial mechanics case.  
Less differentiability for (attracting) charged particles.



Proposition (J. Fejoz, AK, R. Montgomery)

*After simultaneous binary regularization the (incomplete)  $n$ -body flow is continuously differentiable.*

..after additional coordinate change...

## Proposition (J. Fejoz, AK, R. Montgomery)

*After simultaneous binary regularization the (incomplete)  $n$ -body flow is continuously differentiable.*

..after additional coordinate change...

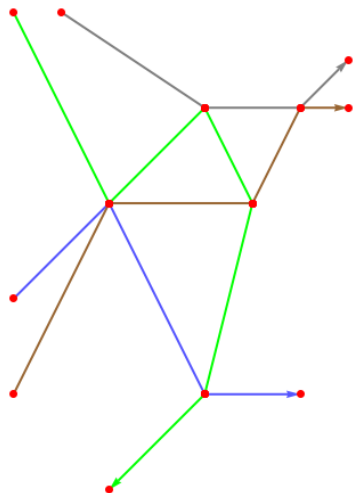
# Train correspondences



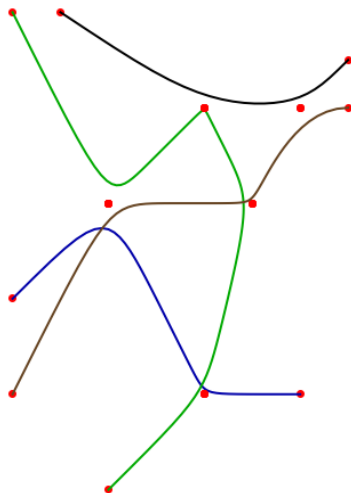
Two Wood Burners in Kamloops Train Yards: Circa 1888

# Specific idea:

find solution where the celestial bodies meet at kinematically prescribed points

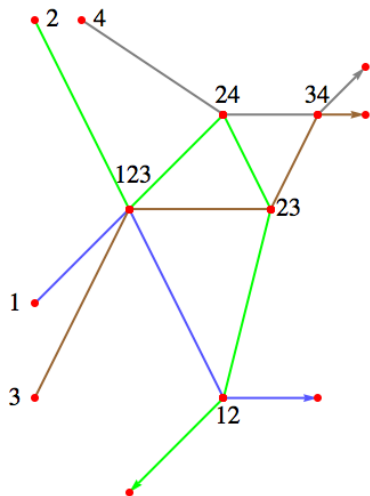


Ideal



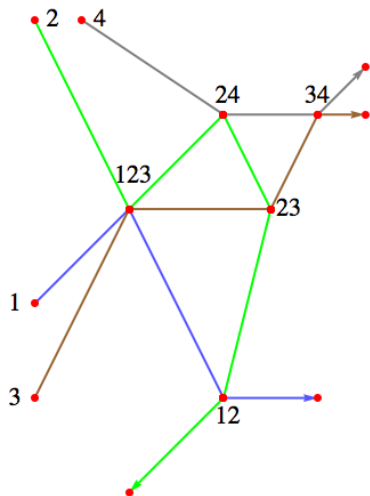
reality (for Coulomb repulsion)

# Kinematics: Setup



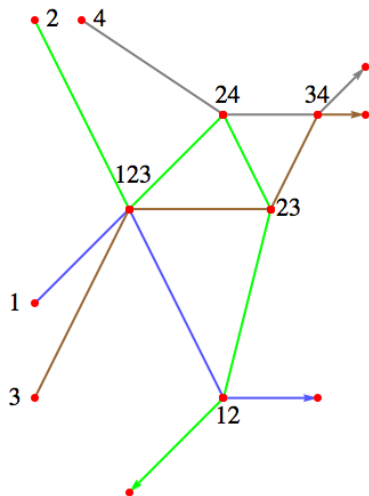
- between collisions, the  $n$  particles (figure:  $n = 4$ ) move with constant velocity
- at collision, there is energy and momentum conservation for each cluster.
- we prescribe the succession of clusters (figure: 123, 24, 23, 34, 12) and consider the variety of all solutions of this non-deterministic dynamics

# Kinematics: Setup



- between collisions, the  $n$  particles (figure:  $n = 4$ ) move with constant velocity
- at collision, there is energy and momentum conservation for each cluster.
- we prescribe the succession of clusters (figure: 123, 24, 23, 34, 12) and consider the variety of all solutions of this non-deterministic dynamics

# Kinematics: Setup

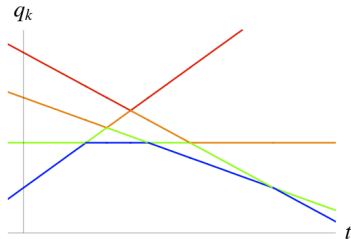
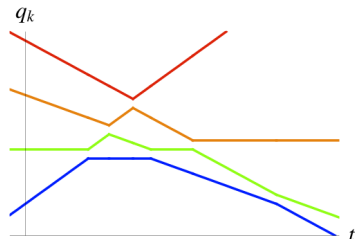


- between collisions, the  $n$  particles (figure:  $n = 4$ ) move with constant velocity
- at collision, there is energy and momentum conservation for each cluster.
- we prescribe the succession of clusters (figure: 123, 24, 23, 34, 12) and consider the variety of all solutions of this non-deterministic dynamics

# Kinematics of $n$ -body scattering

(the hard problems!)

- **Question:** How many collisions for  $n$  balls on the line, having equal mass? (deterministic for binary collisions)



**Answer:**  $\binom{n}{2}$ , look at position  $q_k(t)$  of ball no.  $k$  at time  $t$ !

- **Question:** How many collisions for  $n$  balls on the line, having nearly equal mass (mass ratios in  $[1 - \varepsilon, 1 + \varepsilon]$ )?

**Answers:** 1) Best upper bound is super-exponential in  $n$

D. Burago, S. Ferleger & A. Kononenko, *Uniform estimates on the number of collisions in semi-dispersing billiards*. *Annals of Mathematics* 147, 695–708 (1998)

2) Between  $\binom{n-1}{1}$  and (at least)  $\binom{n+1}{3}$ , depending on precise masses

A. Knauf, M. Stepan: *Elastic Scattering of Point Particles with Nearly Equal Masses* (2019)

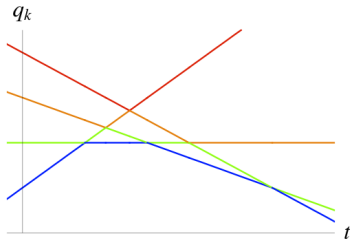
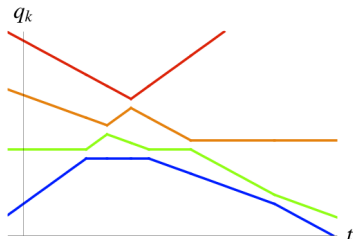




# Kinematics of $n$ -body scattering

(the hard problems!)

- **Question:** How many collisions for  $n$  balls on the line, having equal mass? (deterministic for binary collisions)



**Answer:**  $\binom{n}{2}$ , look at position  $q_k(t)$  of ball no.  $k$  at time  $t$ !

- **Question:** How many collisions for  $n$  balls on the line, having nearly equal mass (mass ratios in  $[1 - \varepsilon, 1 + \varepsilon]$ )?

**Answers:** 1) Best upper bound is super-exponential in  $n$

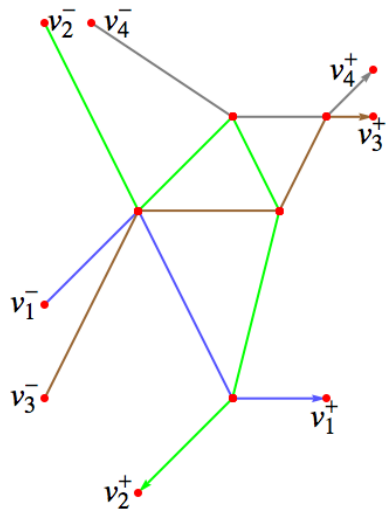
D. Burago, S. Ferleger & A. Kononenko, *Uniform estimates on the number of collisions in semi-dispersing billiards*. *Annals of Mathematics* **147**, 695–708 (1998)

2) Between  $\binom{n-1}{1}$  and (at least)  $\binom{n+1}{3}$ , depending on precise masses

A. Knauf, M. Stepan: Elastic Scattering of Point Particles with Nearly Equal Masses (2010).



# Kinematics: No way from $B$ to $A$ !



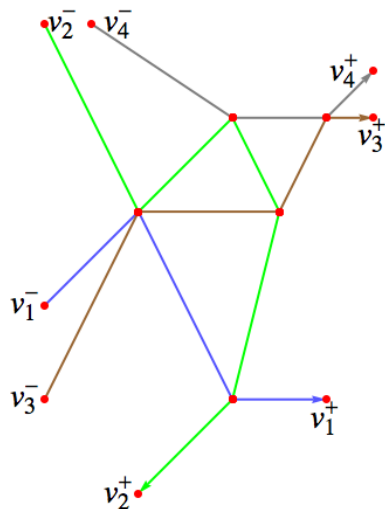
- Consider for given succession of clusters all pairs

$$(v^-, v^+) \in S^D \times S^D$$

of initial/final velocities  $v^\pm$ , with  $D := d(n-1) - 1$  (conservation of total kinetic energy and of total momentum)

- Whereas the  $n$ -body collision has codimension zero,
- for binary clusters the codimension is at least  $n - 2$ .

# Kinematics: No way from $B$ to $A$ !



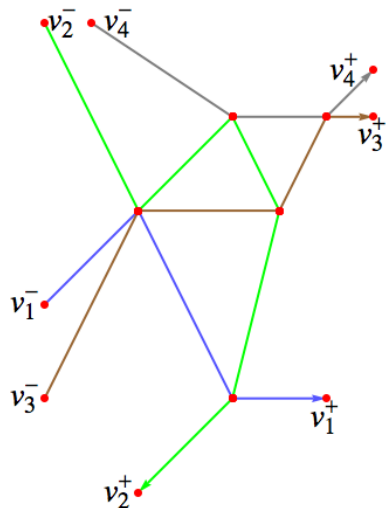
- Consider for given succession of clusters all pairs

$$(v^-, v^+) \in S^D \times S^D$$

of initial/final velocities  $v^\pm$ , with  $D := d(n-1) - 1$  (conservation of total kinetic energy and of total momentum)

- Whereas the  $n$ -body collision has codimension zero,
- for binary clusters the codimension is at least  $n - 2$ .

# Kinematics: No way from $B$ to $A$ !



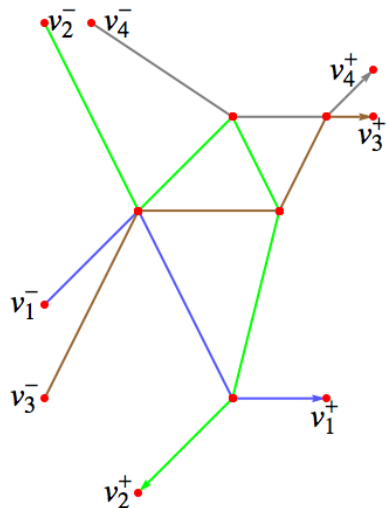
- Consider for given succession of clusters all pairs

$$(v^-, v^+) \in S^D \times S^D$$

of initial/final velocities  $v^\pm$ , with  
 $D := d(n-1) - 1$   
(conservation of total kinetic energy and of total momentum)

- Whereas the  $n$ -body collision has codimension zero,
- for binary clusters the codimension is at least  $n-2$ .

# Kinematics: No way from $B$ to $A$ !



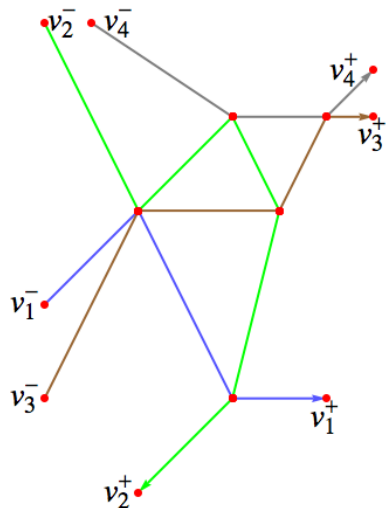
- Consider for given succession of clusters all pairs

$$(v^-, v^+) \in S^D \times S^D$$

of initial/final velocities  $v^\pm$ , with  
 $D := d(n-1) - 1$   
(conservation of total kinetic energy and of total momentum)

- Whereas the  $n$ -body collision has **codimension** zero,
- for binary clusters the codimension is at least  $n-2$ .

# Kinematics: No way from $B$ to $A$ !



- Consider for given succession of clusters all pairs

$$(v^-, v^+) \in S^D \times S^D$$

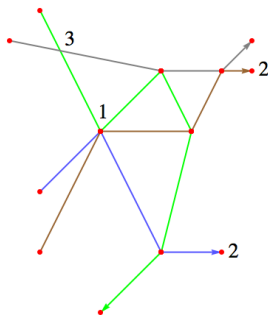
of initial/final velocities  $v^\pm$ , with  
 $D := d(n-1) - 1$   
(conservation of total kinetic energy and of total momentum)

- Whereas the  $n$ -body collision has **codimension** zero,
- for binary clusters the codimension is at least  $n - 2$ .

# Transversality; long range of gravity

**Transversality:** We only consider kinematic solutions for which

- 1 all collisions are binary
- 2 the initial velocities of all particles are different (same for final velocities)
- 3 forward scattering (that is, collision without scattering) does not occur.
- 4 No particle is allowed to go through without any deflection.



**Long range of the  $1/r$  potential:** already Kepler hyperbolae are not asymptotic in time to *any* straight line.

Thus kinematic solutions can at most be approximated by  $n$ -body solutions **locally uniformly** in time.

# Main result

Rescale all interactions:  $I_{k,l} \mapsto \varepsilon I_{k,l}$ .

## Theorem

*In the limit  $\varepsilon \searrow 0$  any transversal kinematic solution is approximated locally uniformly by  $n$ -body solutions.*

## Corollary

*Homogeneity of the kinetic and the potential part in  $H$  leads to corresponding  $n$ -body solutions for any given positive energy, with interaction  $I_{k,l}$  unchanged but spatial scale  $\nearrow \infty$ .*



# Asymptotic velocity

(discussion for smooth long-ranged pair potentials)

$$\bar{v} : T^*\mathbb{R}^{nd} \rightarrow \mathbb{R}^{nd} \quad , \quad \bar{v}(x) := \lim_{t \rightarrow +\infty} \frac{q(t, x)}{t}$$

## Known:

- Asymptotic velocity  $\bar{v}$  **exists**.
- $\bar{v}$  is **discontinuous**.

## Theorem (FKM)

On the *free* phase space region  $P_{free} := \bar{v}^{-1}(\mathbb{R}^{nd} \setminus \Delta)$  it is *smooth*.

# Questions

- 1 Can we obtain a global topological picture of **three-body scattering** in celestial mechanics?
- 2 Can we show for  $m$  electrons in a **molecule** with  $n$  fixed nuclei **existence** of solution  $x : \mathbb{R} \rightarrow P$  for **a.e.** initial condition  $x_0 \in P$ ?

THANK  
YOU